# Tangential Trapezoid Central Configurations 

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#### Abstract

A tangential trapezoid, also called a circumscribed trapezoid, is a trapezoid whose four sides are all tangent to a circle within the trapezoid: the in-circle or inscribed circle. In this paper we classify all planar four-body central configurations, where the four bodies are at the vertices of a tangential trapezoid.


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## 1. INTRODUCTION

The classical $n$-body problem concerns the study of the dynamics of $n$ particles interacting among themselves by their mutual attraction according to Newtonian gravity.

Let $x_{i} \in \mathbb{R}^{d}(i=1, \ldots, n)$ denote the position vector of the $i$-body, and let $m_{i} \in \mathbb{R}^{+}(i=1, \ldots, n)$ denote the mass of the $i$-body. $\mathbb{R}^{d}$ is the Euclidean space ( $d=2$ or 3 ). By Newton's law of motion and Newton's gravitational law the equations of motion of the $n$-body problem are governed by

$$
\ddot{x}_{i}=-\sum_{j=1, j \neq i}^{n} \frac{m_{j}\left(x_{i}-x_{j}\right)}{r_{i j}^{3}}, \quad 1 \leqslant i \leqslant n
$$

where $r_{i j}=\left|x_{i}-x_{j}\right|$ is the mutual Euclidean distance between the $i$-body and the $j$-body. Here we take the gravitational constant $G=1$.

The vector $x=\left(x_{1}, \ldots, x_{n}\right) \in\left(\mathbb{R}^{d}\right)^{n}$ is called the configuration of the system. Define $\delta(x)$ as the dimension of a configuration $x$, i. e., the dimension of the smallest affine space of $\mathbb{R}^{d}$ containing all of the points $x_{i}$. Configurations with $\delta(x)=1,2,3$ are called collinear, planar and spacial, respectively.

When $n=2$, the $n$-body problem has been completely solved. However, for the $n$-body problem for $n \geqslant 3$ the complete solution remains open.

Let

$$
M=m_{1}+\cdots+m_{n}, \quad c=\frac{m_{1} x_{1}+\cdots+m_{n} x_{n}}{M}
$$

be the total mass and the center of masses of the $n$ bodies, respectively.
A configuration $x$ is called a central configuration if the acceleration vectors of the $n$ bodies are proportional to their positions with respect to the center of masses with the same constant of proportionality, i. e.,

$$
\begin{equation*}
\sum_{j=1, j \neq i}^{n} \frac{m_{j}\left(x_{j}-x_{i}\right)}{r_{i j}^{3}}=\lambda\left(x_{i}-c\right), \quad 1 \leqslant i \leqslant n \tag{1.1}
\end{equation*}
$$

where $\lambda$ is the constant of proportionality.

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